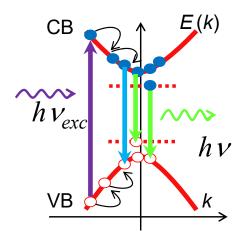
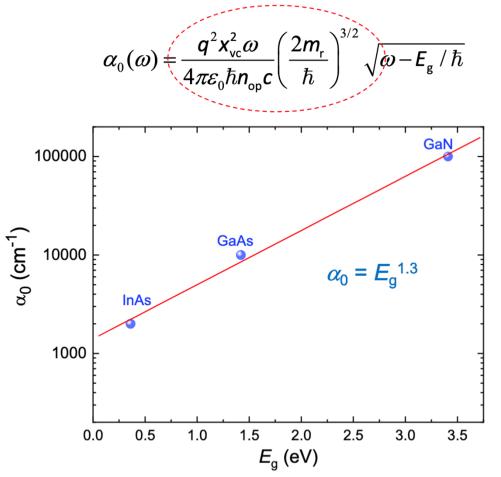
Lecture 14 – 18/12/2024

- Oscillator strength of excitons in semiconductors
- Electron-photon interaction beyond the semi-classical picture
- Spontaneous emission rate and bimolecular recombination coefficient
- Basic insights into near band edge photoluminescence transitions



Summary Lecture 13 - Absorption in semiconductors



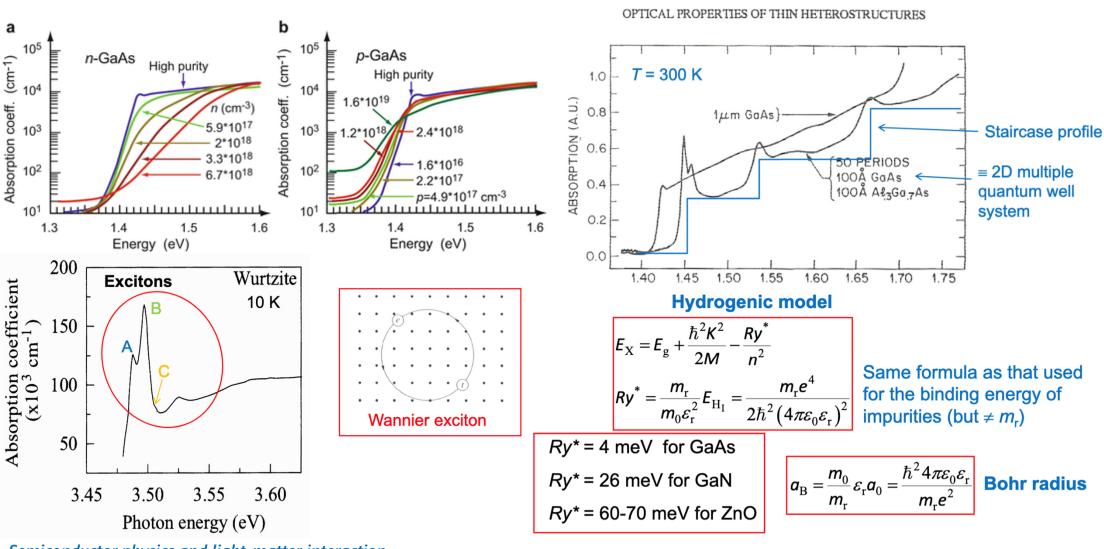
How does the absorption explicitly depend on the bandgap? \rightarrow Missing dependence of n_{op}

$$\Rightarrow P(t) = qx(t)N = \frac{q^2NE(t)}{\left[m_0(\omega_0^2 - \omega^2)\right]} \qquad \bullet \qquad \varepsilon = D/E = \varepsilon_0 + P/E$$

$$\Rightarrow n_{op}^2(\omega) = 1 + \frac{q^2N}{\varepsilon_0 m_0(\omega_0^2 - \omega^2 + i\gamma\omega)}$$

$$\Rightarrow n_{op}(\lambda) = \sqrt{a + \frac{b\lambda^2}{\lambda^2 - c^2}} \qquad \bullet \qquad \bullet \qquad \bullet \qquad \bullet$$
 damping

Summary Lecture 13 - Doping and Excitons



Optical transition probability per unit time for converting a photon into an exciton, $R_{\rm exc}$:

$$R_{\text{exc}} = \frac{2\pi}{\hbar} \sum_{f} \left| \left\langle f \left| \mathcal{H}_{\text{xR}} \right| 0 \right\rangle \right|^{2} \delta \left(\mathcal{E}_{f} \left(\mathbf{K} \right) - \mathcal{E}_{0} - \hbar \omega \right)$$

Fermi's Golden Rule for excitons

Exciton-photon interaction

Wavefunction accounting for the relative motion of e-h pairs (cf. hydrogenic model)

$$\langle f | H_{\mathsf{xR}} | 0 \rangle = \sum_{r,k} \left(1 / \sqrt{N} \right) e^{i\mathbf{k} \cdot \mathbf{r}} \phi_{\mathsf{nlm}}(\mathbf{r}) \langle \psi_{\mathsf{k}}(\mathbf{r}_{\mathsf{e}}) \psi_{-\mathsf{k}}(\mathbf{r}_{\mathsf{h}}) | H_{\mathsf{xR}} | 0 \rangle = \sum_{r,k} \left(1 / \sqrt{N} \right) e^{i\mathbf{k} \cdot \mathbf{r}} \phi_{\mathsf{nlm}}(\mathbf{r}) \langle \psi_{\mathsf{k}}^{\mathsf{c}} | H_{\mathsf{eR}} | \psi_{\mathsf{k}}^{\mathsf{v}} \rangle$$
with

with

Independent of **k**

$$\phi_{nlm}(\mathbf{r}) = R_{nl}(\mathbf{r})Y_{lm}(\theta,\phi)$$
 and $H_{eR} = -e\mathbf{r}.\mathbf{E}$

Electron-radiation interaction Hamiltonian in the electric dipole approximation

Associated Laguerre Spherical harmonic polynomials

$$\sum_{\mathbf{k}} e^{i\mathbf{k}.\mathbf{r}} = N\delta(\mathbf{r}) \Rightarrow \left| \left\langle f \left| H_{\mathsf{xR}} \left| 0 \right\rangle \right|^{2} \right| = N\left| \phi_{\mathsf{nlm}}(0) \right|^{2} \left| \left\langle \psi_{\mathbf{k}}^{\mathsf{c}} \left| H_{\mathsf{eR}} \left| \psi_{\mathbf{k}}^{\mathsf{v}} \right\rangle \right|^{2}$$

Probability of exciting an exciton optically is ∞ overlap of electron-hole wavefunctions

Non-zero only for $l = 0 \Rightarrow$ only excitons with an s symmetry can be optically excited

Power loss of EM field due to absorption in a unit volume of a dielectric medium:¹

Power loss = $R_{\text{oep}} \hbar \omega$ Transition rate in the one-electron picture (oep) Energy lost per $I(z) = I_0 e^{-\alpha z}$ and $\tilde{n} = n_{\text{op}} + j\kappa$

unit of time

$$-\frac{dI}{dt} = -\left(\frac{dI}{dz}\right)\left(\frac{dz}{dt}\right)$$
Phase velocity
$$\Rightarrow \frac{dI}{dz} = -\alpha I_0 e^{-\alpha z} = -\alpha I \text{ and } \frac{dz}{dt} = \begin{bmatrix} c \\ n_{\text{op}} \end{bmatrix} \Rightarrow -\frac{dI}{dt} = \frac{c}{n_{\text{op}}} \alpha I$$

$$\alpha = \frac{4\pi\kappa}{\lambda} \Rightarrow \alpha = \frac{\varepsilon_{\text{i,oep}}\omega}{cn_{\text{op}}} \Rightarrow -\frac{dI}{dt} = \frac{\varepsilon_{\text{i,oep}}\omega I}{n_{\text{op}}^2}$$

$$R_{\text{oep}}\hbar\omega = \frac{\varepsilon_{i,\text{oep}}\omega I}{n_{\text{op}}^{2}} \text{ and per definition } I = \frac{n_{\text{op}}^{2}}{8\pi} |E(\omega)|^{2}$$

$$\Rightarrow \varepsilon_{i,\text{oep}} = \frac{8\pi R_{\text{oep}}\hbar}{|E(\omega)|^{2}}$$

$$R_{\text{oep}} = \frac{2\pi}{\hbar} \sum_{\mathbf{k}, \mathbf{k}, \mathbf{k}} \left| \left\langle \psi_{\mathbf{k}}^{c} \right| H_{\text{eR}} \left| \psi_{\mathbf{k}}^{v} \right\rangle \right|^{2} \delta \left(E_{c} \left(\mathbf{k} \right) - E_{v} \left(\mathbf{k} \right) - \hbar \omega \right),$$

$$\Rightarrow R_{\text{oep}} = \frac{2\pi}{\hbar} \left(\frac{e}{m_0 \omega} \right)^2 \left| \frac{E(\omega)}{2} \right|^2 \sum_{\mathbf{k}} |\mathbf{P}|^2 \delta\left(E_{\text{c}}(\mathbf{k}) - E_{\text{v}}(\mathbf{k}) - \hbar \omega \right)$$
 Absorption transition rate per unit volume of the crystal (See Yu&Cardona for technical details)

Electric dipole transition matrix element in the **p** representation

Imaginary part of the dielectric function for the continuum of states in the one-electron picture $(\varepsilon_{\text{i.oep}})$:¹

$$\varepsilon_{i,\text{oep}} = \frac{1}{4\pi\varepsilon_0} \left(\frac{2\pi e}{m_0 \omega} \right)^2 \sum_{\mathbf{k}} |P|^2 \delta\left(E_c(\mathbf{k}) - E_v(\mathbf{k}) - \hbar\omega \right)$$

Recommended homework:

Show using this expression of $\varepsilon_{\rm i,oep}$ that you can recover the expression of $\alpha_{\rm 0}(\omega)$ obtained in slide 16 of Lecture 12

Imaginary part of the dielectric function for the exciton $(\varepsilon_{i,exc})$:

$$\varepsilon_{\rm i,exc} = \frac{2}{4\pi\varepsilon_0} \left(\frac{2\pi e}{m_0\omega}\right)^2 \sum_{n=1}^{\infty} N \left|\phi_{nlm}\left(0\right)\right|^2 \left|P\right|^2 \delta\left(\hbar\omega - \hbar\omega_n\right) \text{ and } \left|\phi_{n0m}\left(0\right)\right|^2 = \frac{\left(na_{\rm B}\right)^{-3}}{\pi}$$
Spin degeneracy

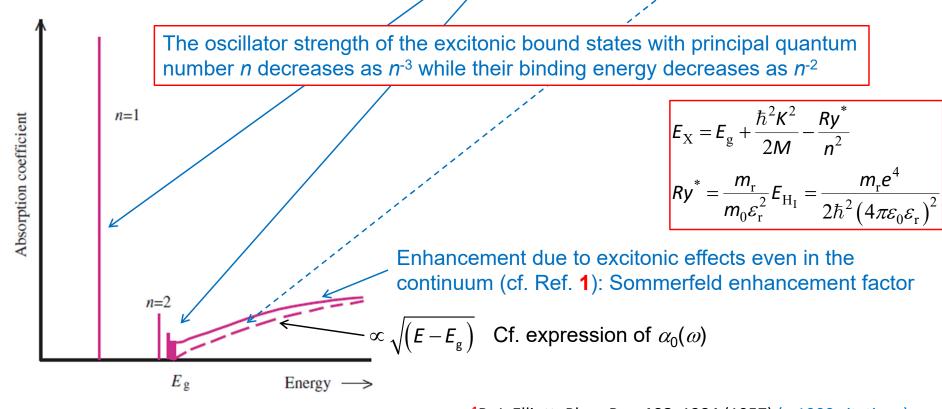
Cf. slide 4

$$\mathcal{E}_{i,\text{exc}}\left(\omega\right) = \frac{N}{32\pi^2} \frac{e^8}{\left(m_0\omega\right)^2} \frac{m_r^3}{\hbar^7 \varepsilon_0^4 \varepsilon_r^3} |P|^2 \sum_{n=1}^{\infty} \frac{1}{n^3} \delta\left(\omega - \omega_n\right)$$
Cf. C. Cohen-Tannoudji, B. Diu, and F. Laloë, Quantum Mechanics, (Wiley-VCH, Weinheim, 2020)

Cf. C. Cohen-Tannoudji, B. Diu, and F. Laloë,

¹Yu & Cardona, Chapter 6

Absorption spectra of a direct gap semiconductor with (solid lines) and without (broken curve) exciton effects for the continuum states^{1,2}



¹R. J. Elliott, Phys. Rev. **108**, 1384 (1957) (> 1900 citations) ²Yu & Cardona, Chapter 6

Free carrier absorption in semiconductors

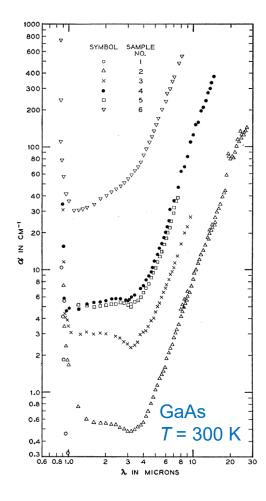


Table I. Electron concentration and doping impurities in GaAs samples used for optical measurements

Sample No.	Donor impurity	n (cm ⁻³)
1		≤5×10 ¹⁴
2	Se	$\overline{1.3} \times 10^{17}$
3		4.9×10^{17}
4	Se	1.09×10^{18}
5	Te	1.12×10^{18}
6	Se	5.4×10^{18}

$$Z_0 \sim 377 \Omega$$
 Impedance of free space

$$\alpha = \frac{\sqrt{\mu_0 / \varepsilon_0} nq^2 \lambda^2}{4\pi^2 c^2 n_{op}} \frac{m^2 \tau^2}{m^* \tau}$$

n: doping level

 λ : wavelength

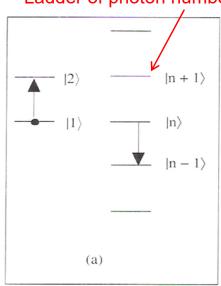
 τ . scattering rate (10⁻¹⁴ s)

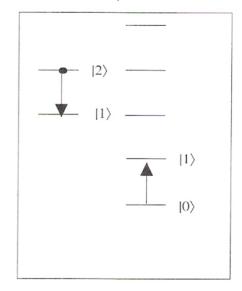
Example

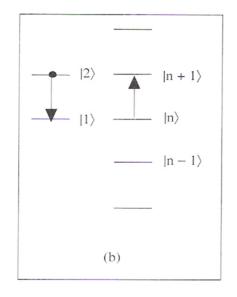
Silicon doped at 10^{20} cm⁻³ $\alpha \approx 2 \times 10^4$ cm⁻¹ at 10 μ m

We consider a 2-level system with states |i, n|, where i is the charged particle state (= 1 or 2 for a 2-level system) and n_l is the number of photons in mode l.

Ladder of photon number states (i.e., Fock states)







(a) Absorption

(b) Spontaneous emission

(c) Stimulated emission

Why does an electron on the excited state relax to the ground state? Or how does it release its energy?

 \Rightarrow Stimulation by the vacuum field fluctuations F_{i}

$$F_{l} = \sqrt{\frac{\hbar \omega_{l}}{2\varepsilon_{0}L^{3}}}$$
 (cavity $V = L^{3}$)

To be admitted

More information in Chaps. 2 & 3

Rosencher-Vinter

The Hamiltonian including the electric field quantization becomes

$$\hat{W} = iqF_{I}(\hat{a}_{I}e^{i\mathbf{k}_{I}\mathbf{r}} - \hat{a}_{I}^{+}e^{-i\mathbf{k}_{I}\mathbf{r}})\mathbf{\varepsilon}_{I}\hat{r}$$

vacuum fluctuations in the mode l

creation and annihilation operators for photons in the mode l

unit polarization vector of the electric field

A few words on the vacuum field fluctuations

Photon number operator in mode I

Hamiltonian of mode l of an EM wave

$$H_{\perp} = \hbar \omega_l \left(\hat{a}_l^{\dagger} \hat{a}_l + \frac{1}{2} \right)$$

Stationary state energy

$$E_{l,i} = \hbar \omega_l \left(i_l + \frac{1}{2} \right)$$
 (Se

Normal component of the Hamiltonian, which is the only one to be of interest (Sec. 2.3 Rosencher-Vinter)

 2 Number of photons in mode $^{\prime}$

 Vacuum energy of mode l (state |0_l>) (inherited from Heisenberg uncertainty principle)

$$E_{vacuum,l} = \frac{1}{2}\hbar\omega_l$$

Mean value and variance of the electric field

$$\frac{\overline{(\hat{\varepsilon}_{\perp l})} = \langle i_l | \hat{\varepsilon}_{\perp l} | i_l \rangle = 0}{\overline{(\hat{\varepsilon}_{\perp l})^2} = \langle i_l | (\hat{\varepsilon}_{\perp l})^2 | i_l \rangle = F_l^2 (2i_l + 1)}$$

with
$$\hat{\varepsilon}_{\perp l} = iF_{l}(\hat{a}_{l}e^{i\mathbf{k}_{l}\mathbf{r}} - \hat{a}_{l}^{+}e^{-i\mathbf{k}_{l}\mathbf{r}})\mathbf{\varepsilon}_{\mathbf{l}}$$

See sections 2.5 and 2.6 in the book by Rosencher & Vinter

Even for an empty cavity (i_l = 0), the variance of the electric field is nonzero and proportional to the square of the *vacuum field fluctuations*

Experimental signature: Lamb shift (degeneracy lift between $2s_{1/2}$ and $2p_{1/2}$ states in the hydrogen atom) signature of the coupling of H atoms with vacuum field fluctuations (full theory: QED theory Feynman-Schwinger-Tomonaga, Nobel Prize 1965)

Let us consider the absorption rate of a 2-level system experiencing an excitation with *n* photons within a mode *l*

The initial state is $|1,n\rangle$ and the final state is $|2,n-1\rangle$: there is 1 photon less

$$P_{1,2} = \frac{2\pi}{\hbar} \left| \langle 2, n-1 | \hat{W} | 1, n \rangle \right|^2 \delta \left(\hbar \omega = E_2 - E_1 \right)$$
 Fermi's Golden Rule

with the field quantization of the perturbation $\hat{W} = iqF_{l}(\hat{a}_{l}e^{i\mathbf{k}_{l}\mathbf{r}} - \hat{a}_{l}^{+}e^{-i\mathbf{k}_{l}\mathbf{r}})\mathbf{\epsilon}_{l}\hat{\mathbf{r}}$

$$P_{1,2} = \frac{2\pi}{\hbar} q^2 F_i^2 \left| \left\langle 2, n - 1 \right| \left(\hat{a}_i e^{i \mathbf{k}_i \mathbf{r}} - \hat{a}_i^+ e^{-i \mathbf{k}_i \mathbf{r}} \right) \mathbf{\epsilon}_1 \hat{\mathbf{r}} \left| 1, n \right\rangle \right|^2 \delta \left(\hbar \omega = E_2 - E_1 \right)$$

The electron state is independent of the photon state (i.e., we have two separable Hilbert spaces)

$$\Rightarrow P_{1,2} = \frac{2\pi}{\hbar} q^2 F_i^2 \left| \langle n - 1 | (\hat{a}_i e^{i\mathbf{k}_i \mathbf{r}} - \hat{a}_i^+ e^{-i\mathbf{k}_i \mathbf{r}}) | n \rangle \right|^2 \left| \langle 2 | \mathbf{\epsilon}_1 \hat{\mathbf{r}} | 1 \rangle \right|^2 \delta(\hbar \omega = E_2 - E_1)$$

$$\langle n - 1 | \hat{a}_i | n \rangle = \sqrt{n} \quad \text{only non-zero related term}$$

The absorption rate is then given by

$$P_{1,2} = \frac{2\pi}{\hbar} q^2 F_1^2 n \left| \langle 2 | \mathbf{\epsilon}_1 \hat{\mathbf{r}} | 1 \rangle \right|^2 \delta(\hbar \omega = E_2 - E_1)$$

$$F_{l} = \sqrt{\frac{\hbar \omega_{l}}{2\varepsilon_{0} L^{3}}}$$

$$P_{1,2} = \frac{2\pi}{\hbar} q^2 \frac{\hbar \omega_l}{2\varepsilon_0 L^3} n \left| \left\langle 2 \left| \mathbf{\varepsilon_l} \hat{\mathbf{r}} \right| 1 \right\rangle \right|^2 \delta(\hbar \omega = E_2 - E_1)$$

On the other hand, the amplitude of the classical electric field is such that $E_0 = \sqrt{\frac{2\hbar\omega_l}{\varepsilon_0 L^3}} n$

$$\boldsymbol{E}_0 = \sqrt{\frac{2\hbar\omega_l}{\varepsilon_0 L^3}} \, \boldsymbol{n}$$

Thus,
$$P_{1,2} = \frac{\pi}{2\hbar} q^2 E_0^2 \left| \left\langle 2 \left| \mathbf{\epsilon_l} \hat{\mathbf{r}} \right| 1 \right\rangle \right|^2 \delta(\hbar \omega = E_2 - E_1)$$
 Cf. Lecture 12, slides 5-8

Absorption rate of a photon (with *n* photons in mode *l*) with an electron transition from state 1 to state 2

The recombination rate is calculated in the same way

The initial state is $|2,n\rangle$ and the final state is $|1,n+1\rangle$: 1 photon is created

$$P_{2,1} = \frac{2\pi}{\hbar} q^{2} F_{l}^{2} \left| \langle n+1 | \left(\hat{a}_{l} e^{i\mathbf{k}_{l}\mathbf{r}} - \hat{a}_{l}^{+} e^{-i\mathbf{k}_{l}\mathbf{r}} \right) | n \rangle \right|^{2} \left| \langle 2 | \mathbf{\epsilon}_{1} \hat{\mathbf{r}} | 1 \rangle \right|^{2} \delta \left(\hbar \omega = E_{2} - E_{1} \right)$$

$$\left\langle n+1 | \hat{a}_{l}^{+} | n \rangle = \sqrt{n+1} \quad only \ non-zero \ related \ term$$

Thus, we have

$$P_{2,1} = \frac{2\pi}{\hbar} q^2 \frac{\hbar \omega_l}{2\varepsilon_0 L^3} (n+1) |\langle 2| \varepsilon_1 \hat{\mathbf{r}} |1 \rangle|^2 \delta(\hbar \omega = E_2 - E_1)$$

Recombination rate of an electron from state 2 to state 1 with 1 photon emitted in mode l

Recombination rate in a two-level system

$$P_{2,1} = q^{2} \frac{\pi \omega_{l}}{\varepsilon_{0} L^{3}} | \langle 2 | \mathbf{\epsilon}_{1} \hat{\mathbf{r}} | 1 \rangle |^{2} \delta(\hbar \omega = E_{2} - E_{1}) + q^{2} \frac{\pi \omega_{l}}{\varepsilon_{0} L^{3}} | \langle 2 | \mathbf{\epsilon}_{1} \hat{\mathbf{r}} | 1 \rangle |^{2} \delta(\hbar \omega = E_{2} - E_{1})$$

Stimulated emission Spontaneous emission (P_{sp})

Stimulated emission \Rightarrow proportional to the photon number n

Spontaneous emission \Rightarrow due to vacuum fluctuations F_{i}

The spontaneous emission rate is calculated over all the cavity modes

$$\Gamma_{\rm sp} = \iiint P_{\rm sp} {\rm d}^3 {\bf k} \qquad \Gamma_{\rm sp} = \frac{{\bf q}^2 r_{12}^2 \omega^3 n_{\rm op}}{3\pi c^3 \hbar \varepsilon_0} = 1/\tau_{\rm sp} \qquad \tau_{\rm sp} \quad {\it radiative lifetime}$$
 + substitute $\varepsilon_0 \to \varepsilon_{\rm r} \varepsilon_0$

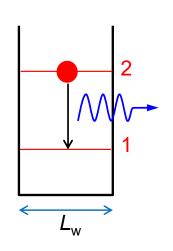
First derivation: Weisskopf-Wigner theory 1930

V. Weisskopf and E. Wigner, Z. Phys. **63**, 54 (1930) (> 1500 citations)

Spontaneous emission: case of discrete levels

Application:

Recombination lifetime in a quantum well with infinite barriers



$$E_n = n^2 \frac{\hbar^2 \pi^2}{2m^* L_w^2}$$



Here, n is the principal quantum number of the quantized energy levels and not the photon number in the optical mode!

Then, $\hbar\omega_{12} = E_2 - E_1 = 3\frac{\hbar^2\pi^2}{2m^*L_w^2}$

$$\hbar\omega_{12} = E_2 - E_1 = 3\frac{\hbar^2\pi^2}{2m^*L_w^2}$$

2) Dipole element r_{12}

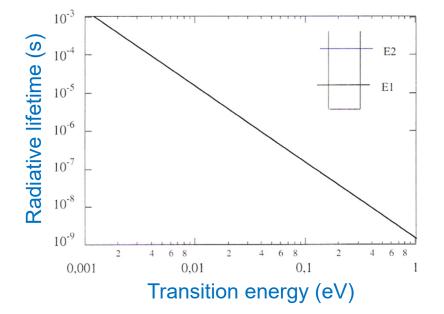
$$r_{12} = \left\langle \Psi_2(x) \middle| x \middle| \Psi_1(x) \right\rangle \quad \text{with} \quad \Psi_1(x) = \sqrt{\frac{2}{L_\text{w}}} \cos(\pi x / L_\text{w}) \quad \text{and} \quad \Psi_2(x) = \sqrt{\frac{2}{L_\text{w}}} \sin(2\pi x / L_\text{w})$$

Finally,
$$r_{12} = \frac{2^4 L_{\text{w}}}{3^2 \pi^2}$$
 and $r_{12}^2 = \frac{2^7 m^*}{3^3 \pi^2 \hbar^2} 1 / E_{12}$

Spontaneous emission: case of discrete levels

$$1/\tau_{\rm sp} = \frac{q^2 r_{12}^2 \omega^3 n_{\rm op}}{3\pi c^3 \hbar \varepsilon_0}$$

$$\tau_{\rm sp} = \frac{3^4 \pi^3 m^* c^3 \hbar^2 \varepsilon_0}{2^7 q^2 n_{\rm op}} \frac{1}{E_{12}^2}$$



Transition energies > 1 eV
⇒ radiative lifetime ~1 ns

Values to be also compared to the optical phonon lifetime (slide #27 of Lecture 6) for a rough comparison!

Spontaneous emission in a bulk system

In an intrinsic bulk semiconductor:

The spontaneous recombination rate $r_{\rm sp}(\mathbf{k})$ (s⁻¹) between the CB and the VB is given for a state with a wavevector **k**

$$r_{\rm sp}(\mathbf{k}) = A_{\rm cv} f_{\rm c}(E_{\rm c}) \left(1 - f_{\rm v}(E_{\rm v})\right)$$

with $A_{CV} = 1/\tau_{R}$ the spontaneous recombination rate

and the radiative lifetime is given by

$$\tau_{\rm R} = \frac{\pi c^3 \hbar \varepsilon_0}{q^2 x_{\rm vc}^2 n_{\rm op} \omega_{\rm vc}^3} = \frac{2\pi c^3 \hbar^2 \varepsilon_0 m_0}{q^2 n_{\rm op} E_{\rm g} E_{\rm p}}$$

$$\tau_{\rm R} \uparrow \text{ when } E_{\rm g} \downarrow \text{ because, overall, the } n_{\rm op} E_{\rm g} E_{\rm p} \text{ product decreases!}$$

⇒ It is more challenging to achieve a laser based on a wide bandgap SC!

Spontaneous emission in a bulk system

In an intrinsic bulk semiconductor:

The spectral distribution of spontaneous recombination rate $R_{\rm sp}(h\nu)$ due to a quasi-equilibrium distribution of carriers is then given by

$$R_{\rm sp}(h\nu) = 2\sum_{\mathbf{k}} r_{\rm sp}(\mathbf{k}) = 2\sum_{\mathbf{k}} \frac{1}{\tau_{\rm R}(\mathbf{k})} f_{\rm c}(\mathbf{k}) (1 - f_{\rm v}(\mathbf{k})) \delta(E_{\rm c} - E_{\rm v} = h\nu)$$

The summation is performed over all the **k**-vectors verifying the energy conservation condition (hence the Dirac delta)

$$E_{c}(\mathbf{k}) - E_{v}(\mathbf{k}) = hv = E_{g} + \frac{\hbar^{2}k^{2}}{2m_{r}}$$

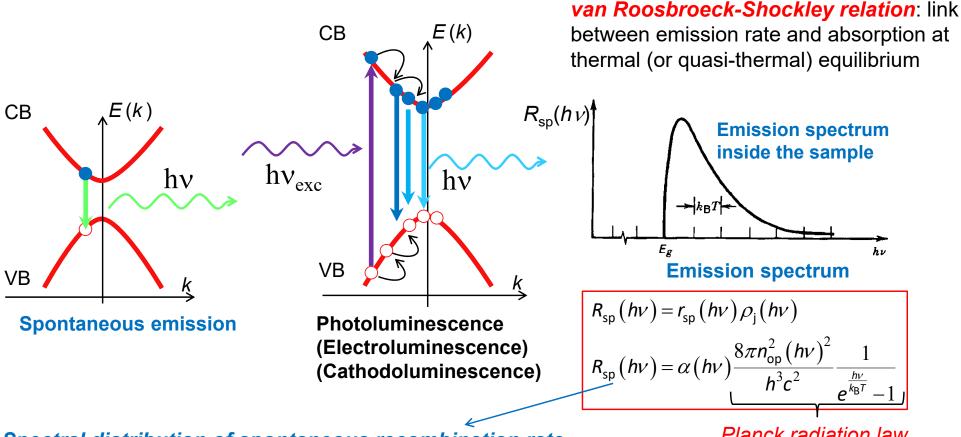
which leads to

$$R_{\rm sp}\left(h\nu\right) = \int_0^\infty r_{\rm sp}\left(E\right) \rho_{\rm j}\left(E\right) \delta\left(E = h\nu\right) {\rm d}E = r_{\rm sp}\left(h\nu\right) \rho_{\rm j}\left(h\nu\right) \quad \text{using the fact that } 2\sum_{\bf k} \leftrightarrow \int_{\bf k} \rho({\bf k}) d^3{\bf k} \leftrightarrow \int_E \rho(E) dE$$

$$R_{\rm sp}(h\nu) = \frac{1}{\tau_{\rm R}} \rho_{\rm j}(h\nu) f_{\rm c}(E_{\rm c}(h\nu)) (1 - f_{\rm v}(E_{\rm v}(h\nu)))$$

Spontaneous emission in a bulk system

Recombination and spontaneous emission



Spectral distribution of spontaneous recombination rate

Planck radiation law

W. van Roosbroeck and W. Shockley, Phys. Rev. 94, 1558 (1954) (> 900 citations)

Bimolecular recombination coefficient

Total radiative recombination rate

$$R_{\rm sp} = \frac{1}{\tau_{\rm R}} \int_{E_{\rm g}}^{\infty} \rho_{\rm j} (h\nu) f_{\rm c} (E_{\rm c} (h\nu)) (1 - f_{\rm v} (E_{\rm v} (h\nu))) dh\nu$$

i.e.,
$$R_{\rm sp} = \frac{e^{\left(E_{\rm F_c} - E_{\rm F_v}\right)/k_{\rm B}T}}{\tau_{\rm R}} \int_{E_{\rm g}}^{\infty} \rho_{\rm j} (h\nu) e^{-h\nu/k_{\rm B}T} {\rm d}h\nu$$
 Expression valid within Boltzmann approximation!

Can be recast into a constant × the gamma function (aka the Euler integral of the 2nd kind)!

which leads to

 $R_{\rm sp} = \frac{1}{\pi} N_{\rm j} e^{(E_{\rm fc} - E_{\rm fv} - E_{\rm g})/k_{\rm B}T}$ with $N_{\rm j}$ the effective density of states with reduced mass $m_{\rm r}$

Finally, we get:
$$R_{\rm sp} = \frac{1}{\tau_{\rm R}} \frac{N_{\rm j}}{N_{\rm c} N_{\rm v}} np$$

with
$$B = \frac{1}{\tau_{\rm R}} \frac{N_{\rm j}}{N_{\rm c} N_{\rm v}} = \frac{1}{\tau_{\rm R} N_{\rm c}} \left(\frac{m_{\rm r}}{m_{\rm v}}\right)^{3/2}$$

$$Cf. \ Lecture \ 7$$

$$Band-to-band$$

$$recombinations$$

$$Slides \ 18-21$$

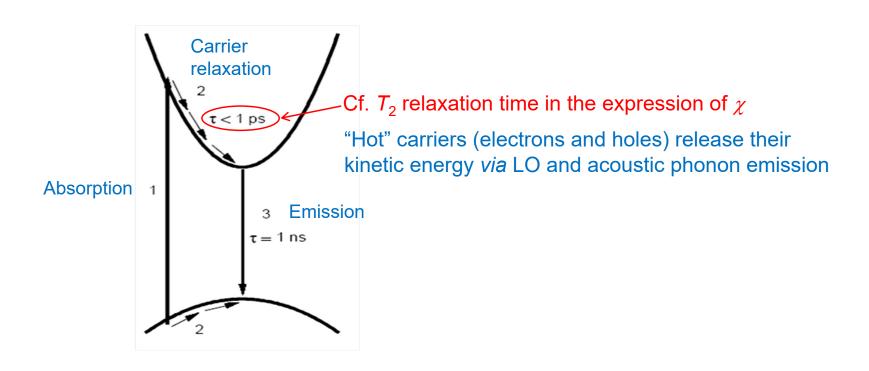
Cf. Lecture 7

Material	<i>B</i> (cm ³ s ⁻¹)
GaAs	7.2×10^{-10}
GaSb	2.4×10^{-10}
InP	1.3 × 10 ⁻⁹
InAs	8.5×10^{-11}
InSb	4.6×10^{-11}

@ 300 K

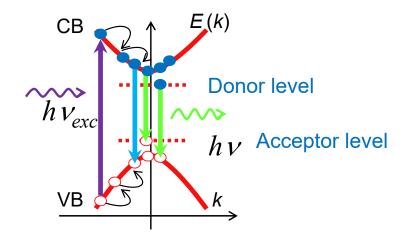
Spontaneous emission

Photoluminescence

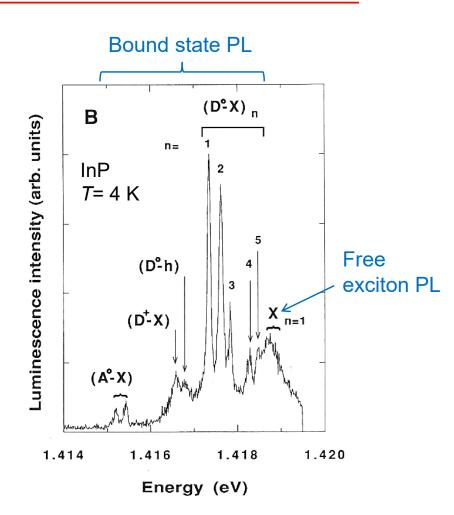


Spontaneous emission

Photoluminescence (case of bulk InP)

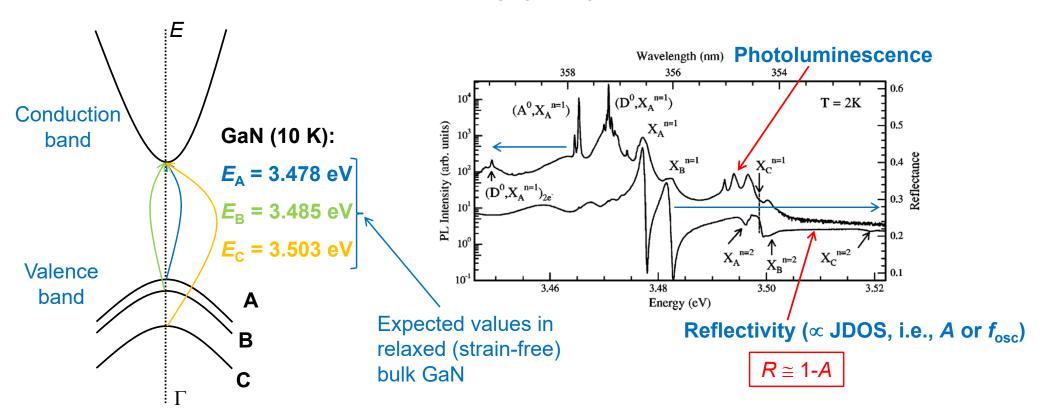


⇒ How can we discriminate intrinsic PL from extrinsic one?



Spontaneous emission

Photoluminescence and reflectivity (GaN)



Exciton binding energy can also be determined from reflectivity spectra!

Examination protocol

Examination: 23/01/2025 from 3:15pm until 6:15pm in room BC 0 1

- The exam could cover any of the topics addressed during the 14 lectures of this semester and the corresponding series.
- The exam will be a mixture of problem solving (involving basic mathematical calculations) and analysis of figures with a main focus on the underlying physics. Special care will be paid to the quality and the correctness of your explanations and/or arguments.
- The exam shall be written in readable English (pay a special attention to your handwriting...).
- Printed versions of the lectures with your personal notes as well as the exercises (+ solutions) are allowed (+ summary of the core concepts posted on Moodle) but you cannot bring anything else (no textbook, etc.)! Any tool providing access to the Internet is strictly forbidden (no smartphone and smartwatch).
- The reference for all the courses of this semester will be the version of the lectures posted on the Moodle repository, i.e., the files available <u>from December 20 2024!</u>
- Do not forget to bring a scientific calculator for numerical applications. There won't be any of them on loan!
- You can reach me preferably by e-mail (from January 3 2025): raphael.butte@epfl.ch (Samuele will always forward your requests to me so please use the shortest pathway!)